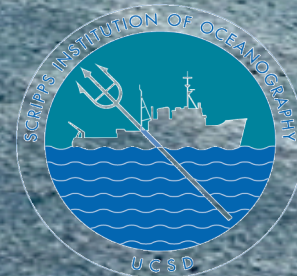


Challenges in measuring near-surface light fields

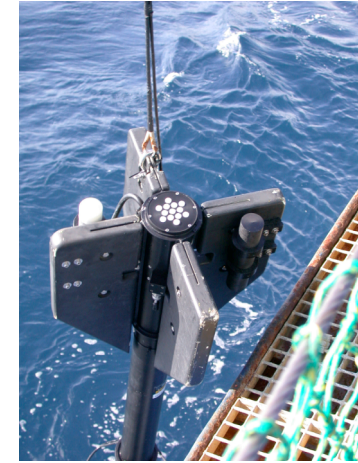
Dariusz Stramski

NASA AOP Workshop
Santa Barbara, CA
13-15 January 2009



Satlantic SPMR in freefall mode

Data recorded at 6 Hz, ~0.14 m resolution
1 deep + 2-4 shallow casts per station
Aim for stable sky conditions



Satlantic ProSoft

Raw data converted to engineering units using calibration file and measured dark currents.
Measured pressure offsets applied.
Tilt edited (data $\geq 5^\circ$ rejected).
 E_d and L_u data interpolated to common depth field (0.1 m).

Custom code

K_{Ed} , K_{Lu} , $E_d(0^-)$, $L_u(0^-)$ calculated from linear LSQ fit of ln-transformed data.

Null depth estimates transferred across air-sea interface:

$$E_d(0^+) = E_d(0^-) / (1 - \alpha) \quad \alpha = 0.043$$

$$L_w = L_u(0^-) [1 - (1 - n_w)^2 / (1 + n_w)^2] n_w^{-2} \quad n_w = 1.343$$

R_{rs} calculated as $L_w / E_d(0^+)$

Averaging of replicates and creation of final data products.

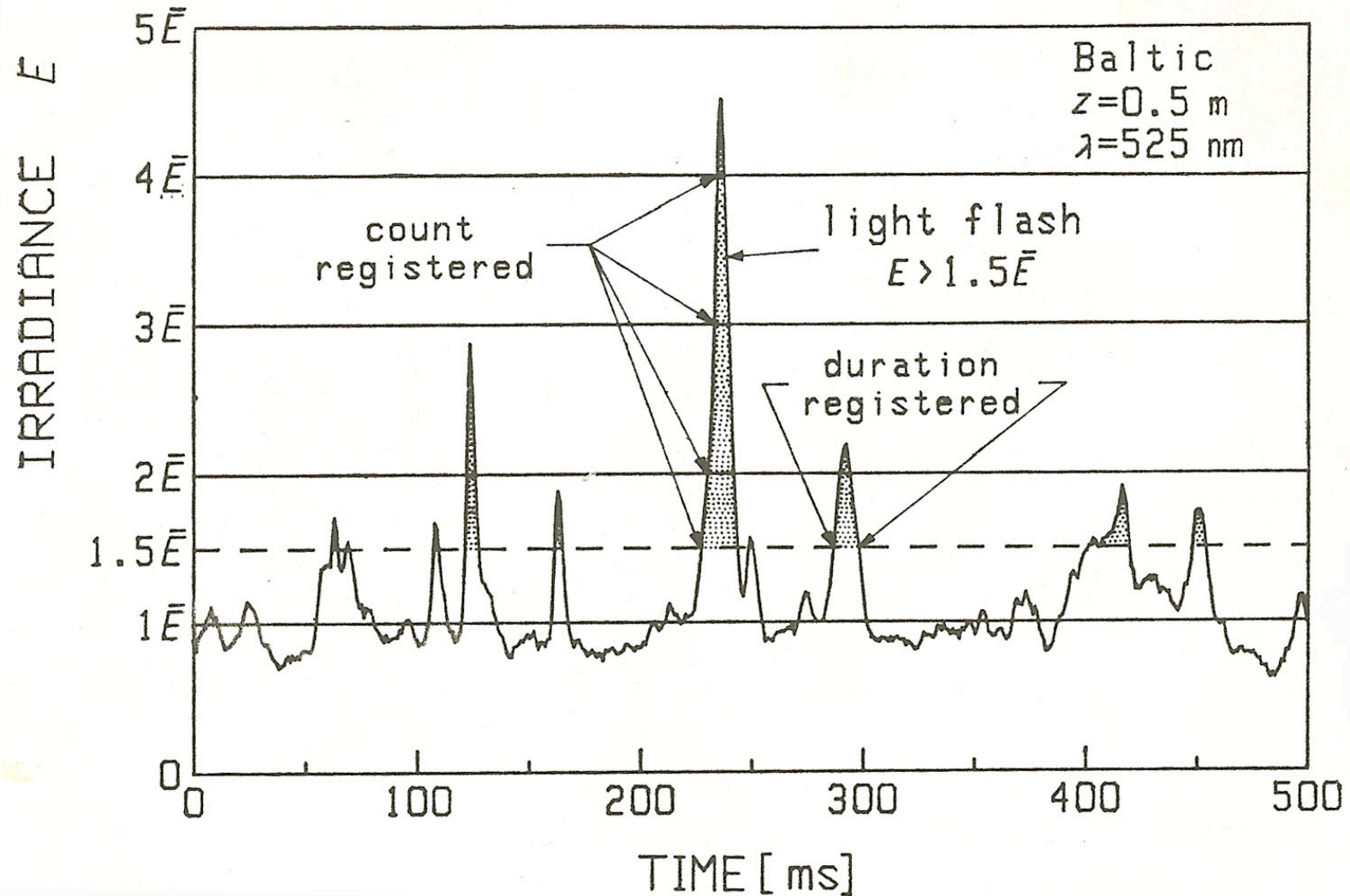
 User Intervention

Visual inspection of individual profiles.
Surface extrapolation interval chosen.

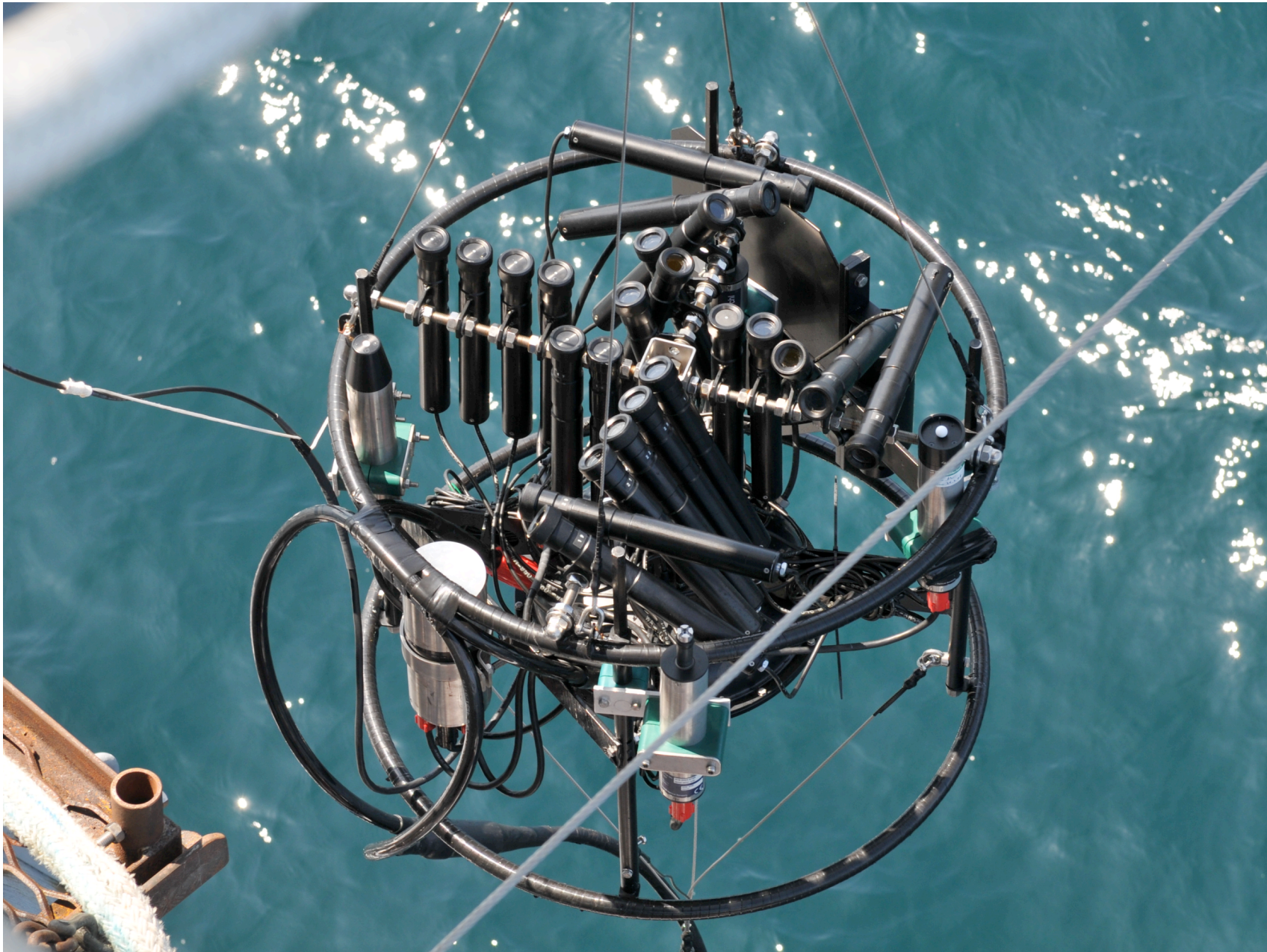


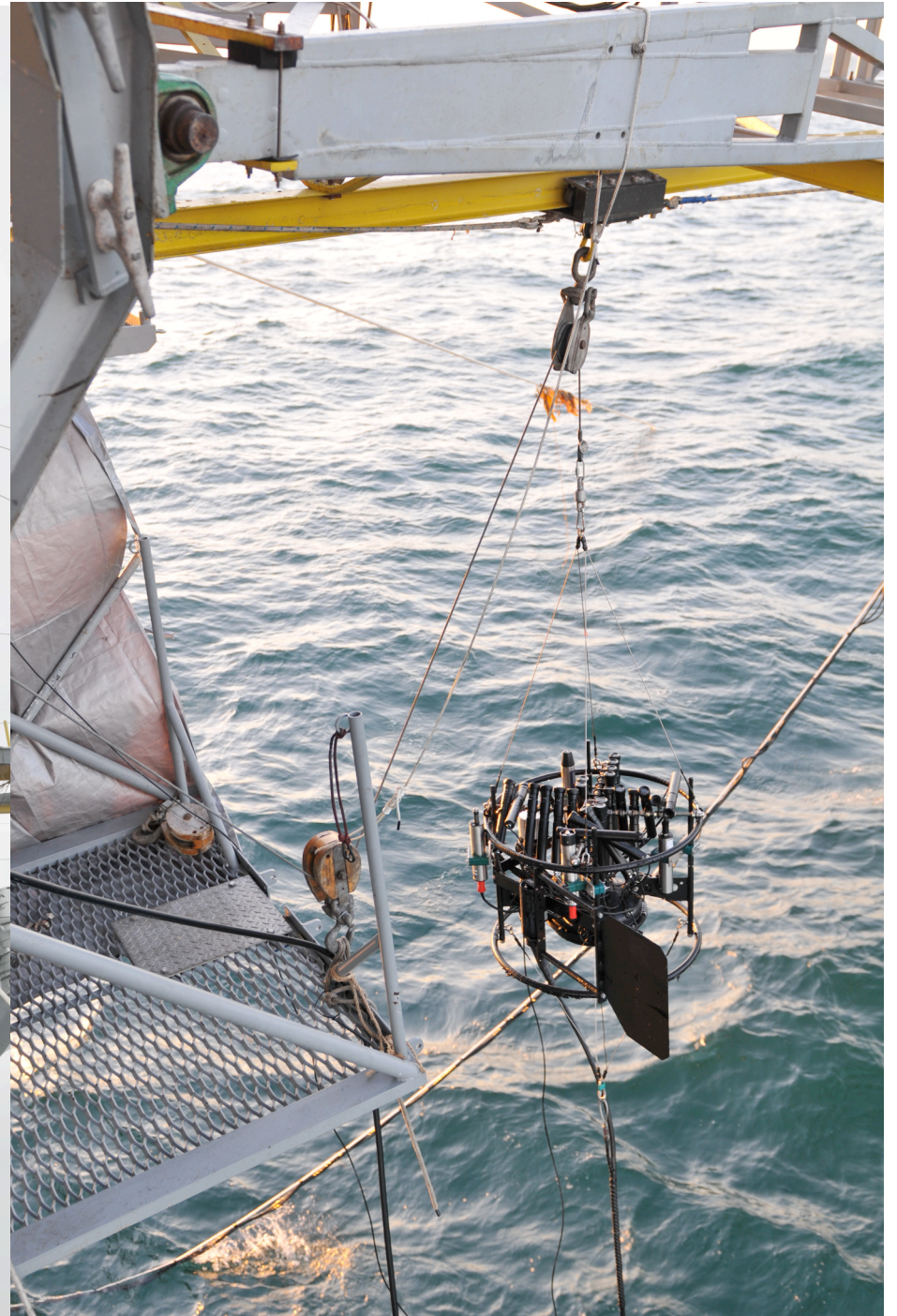
Visual inspection for quality control and reproducibility

Basic principles involved in threshold analysis of light fluctuations (Dera and Stramski 1986)



Underwater Porcupine Radiometer System

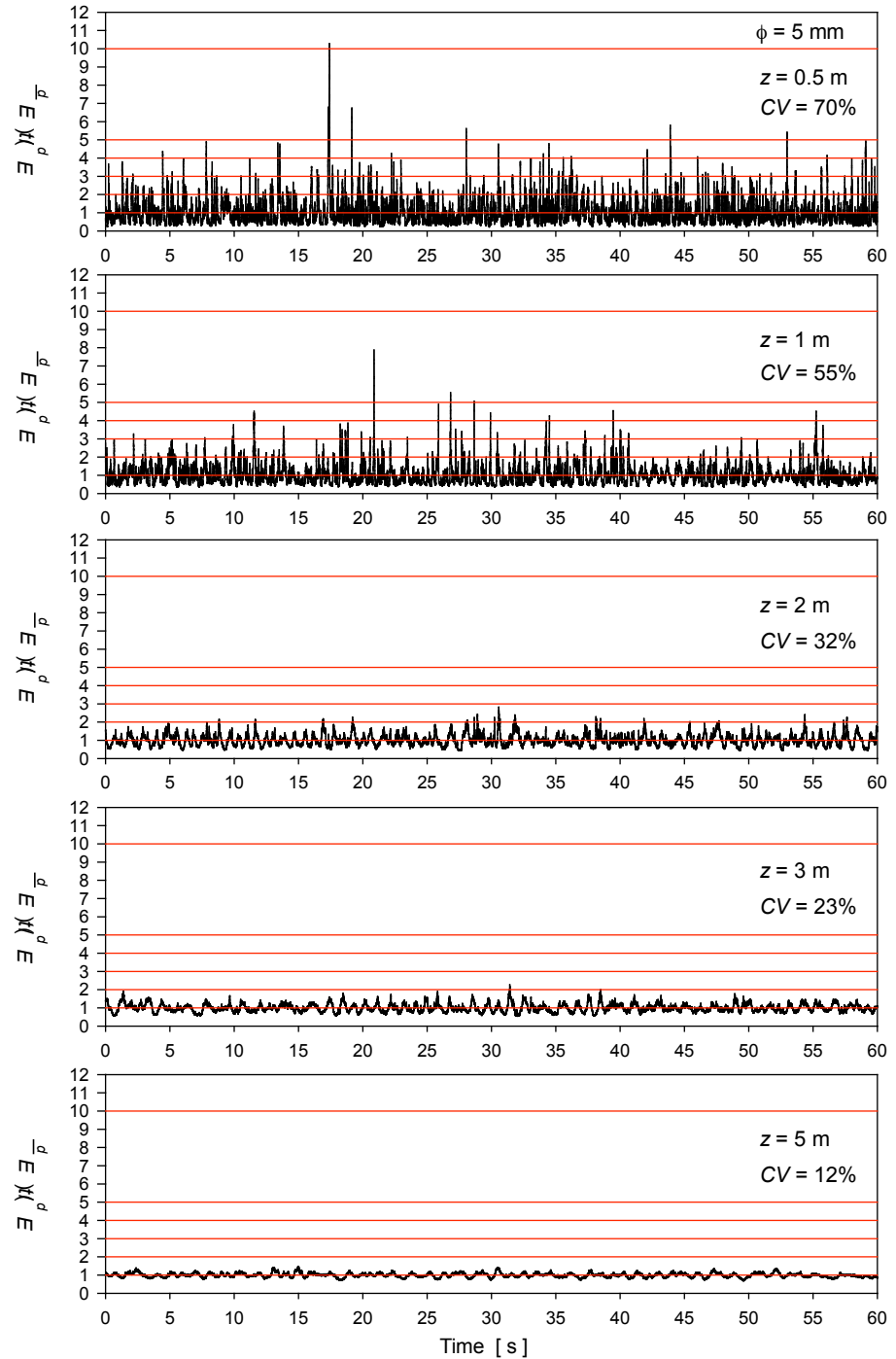




Scripps Pier, January 22, 2008

Changes in light fluctuations with depth

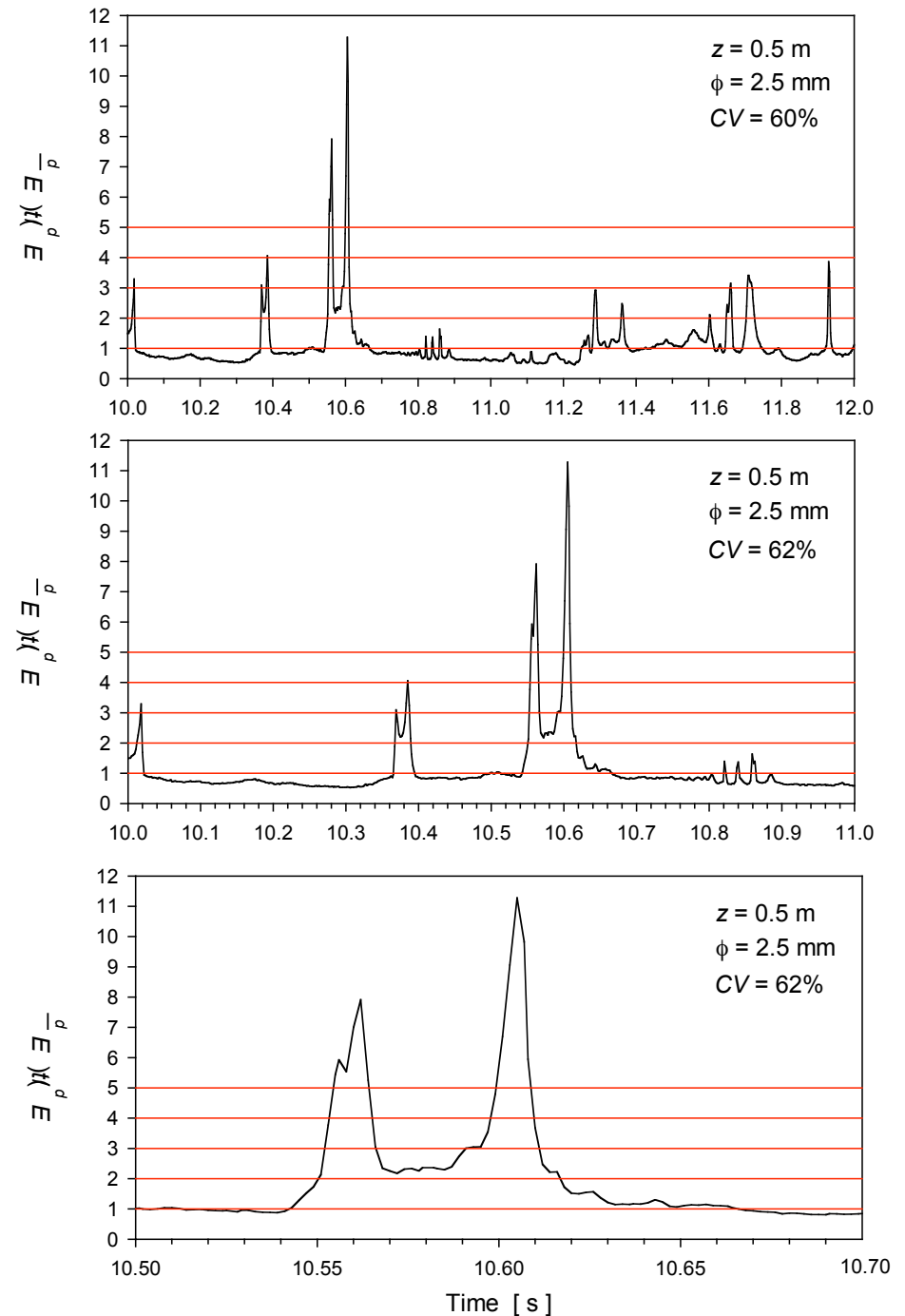
Example 1-min time-series of normalized downwelling irradiance at different depths



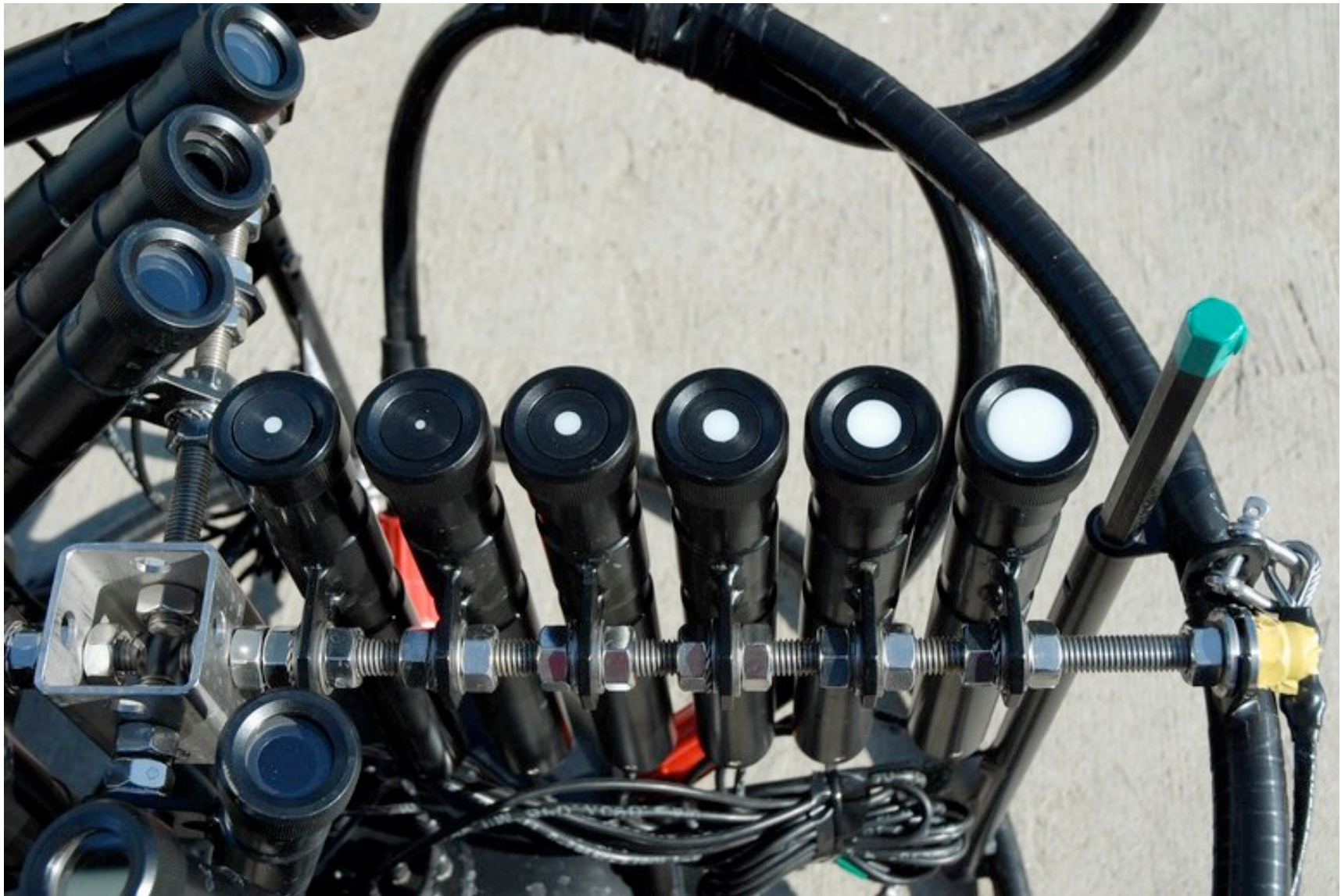
Detailed view of individual wave-focusing events that show very strong pulses of short duration in the E_d signal

The most intense pulses of focused light exceed the time-averaged irradiance by a factor greater than 10.

The duration of pulses is on the order of milliseconds to tens of milliseconds.

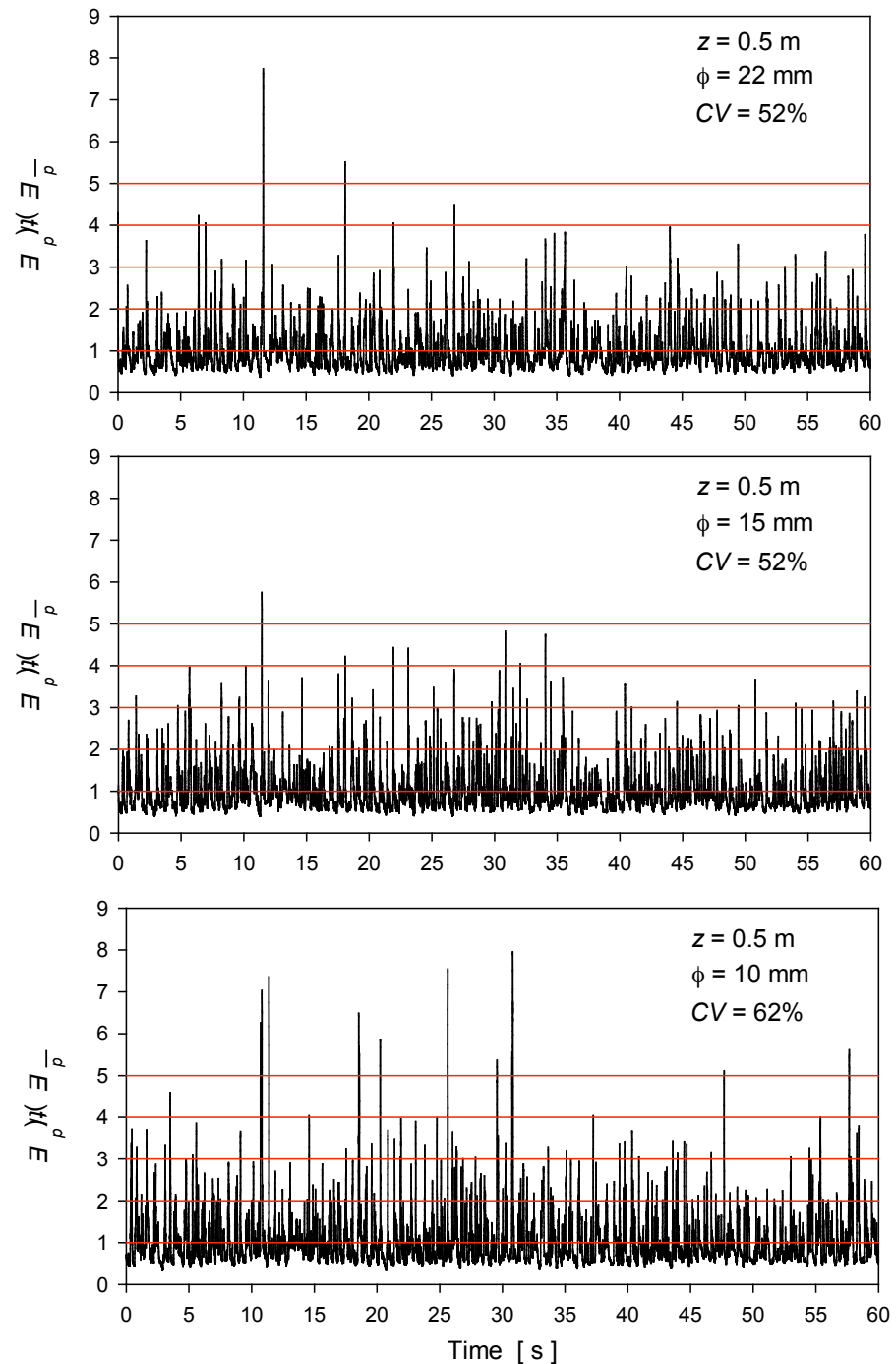


E_d sensors with different surface area of the cosine collector

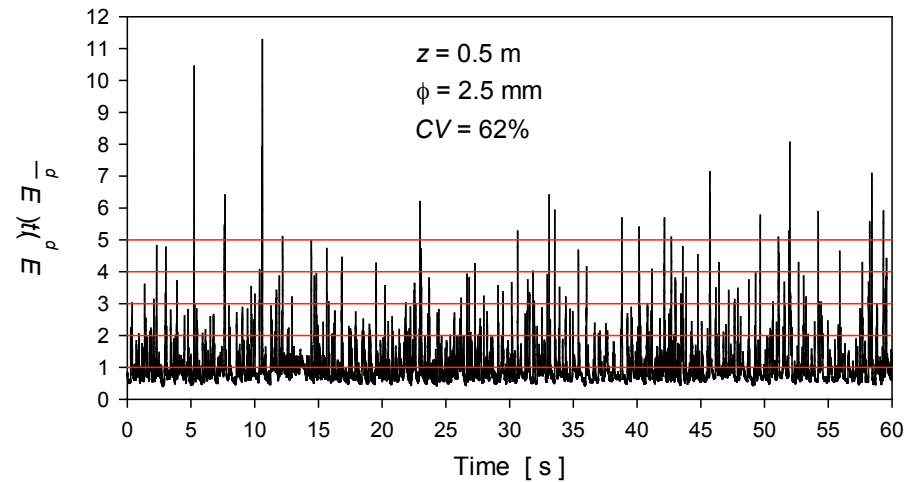
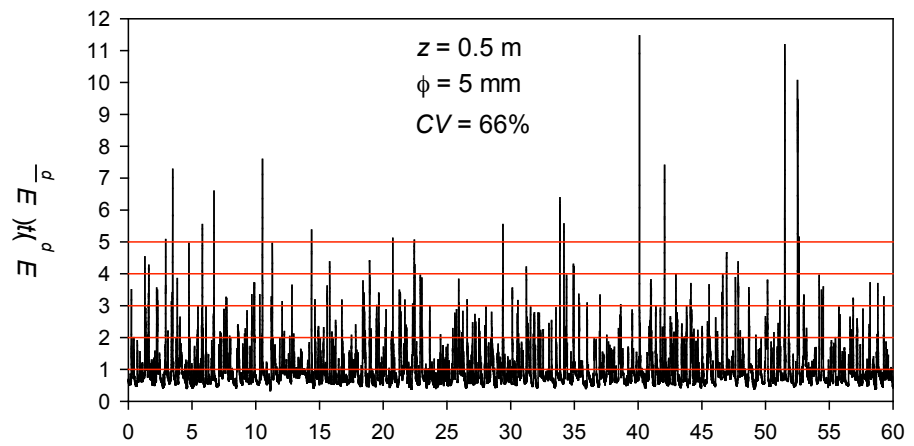
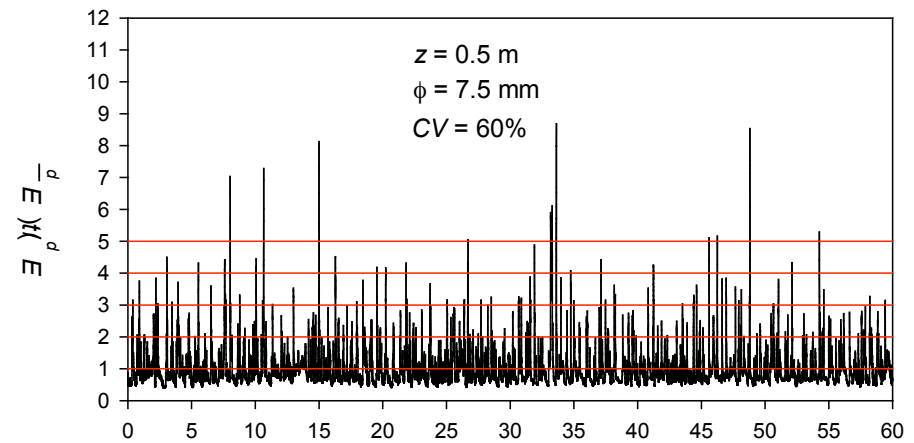


Scripps Pier, Jan 2008

Example time-series of $E_d(532 \text{ nm})$ at $z = 0.5 \text{ m}$ obtained with cosine collectors of different surface area



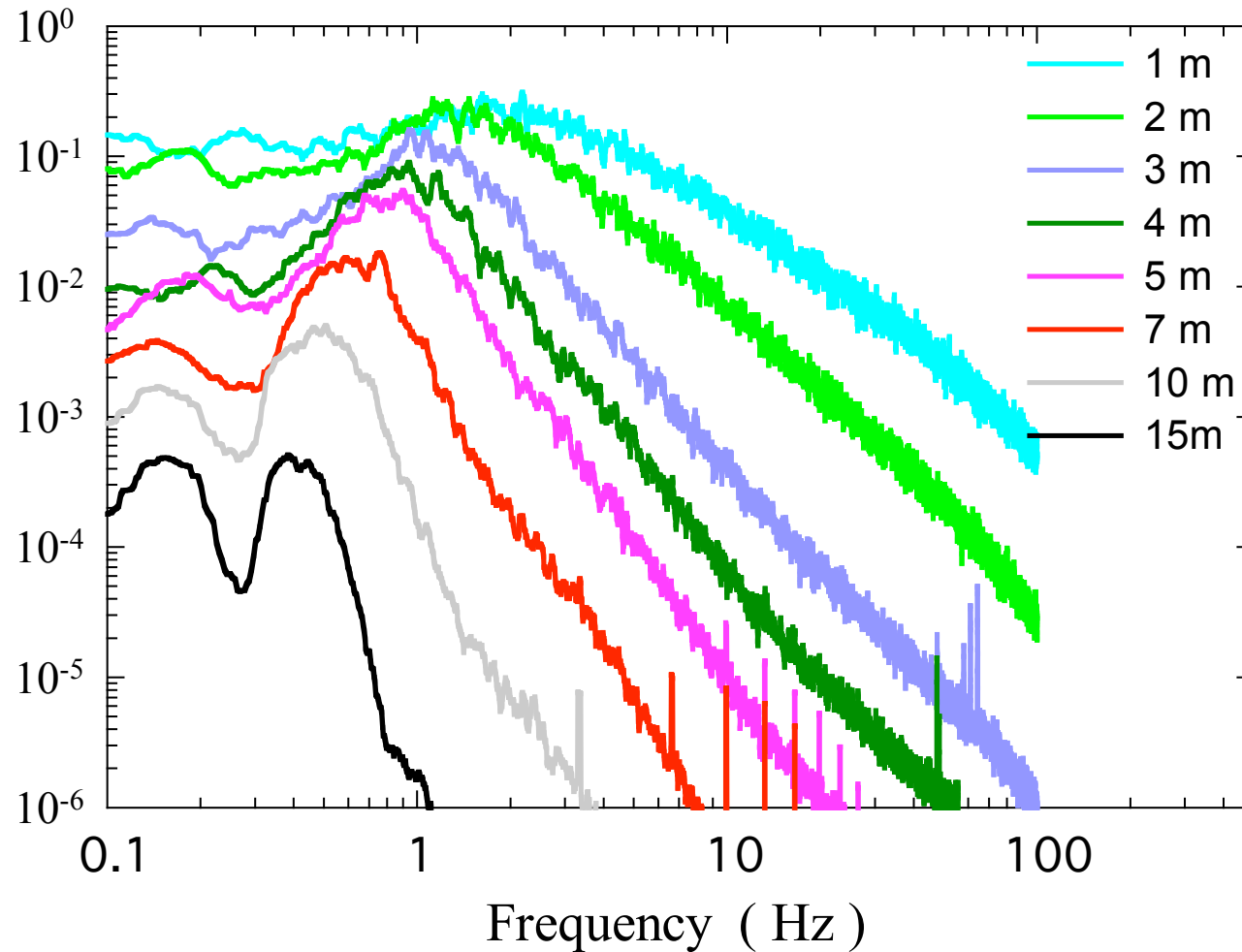
Example time-series of $E_d(532 \text{ nm})$ at $z = 0.5 \text{ m}$ obtained with cosine collectors having different surface area (contd.)



Power spectra of downward irradiance fluctuations

Santa Barbara Channel, 11 September 2008

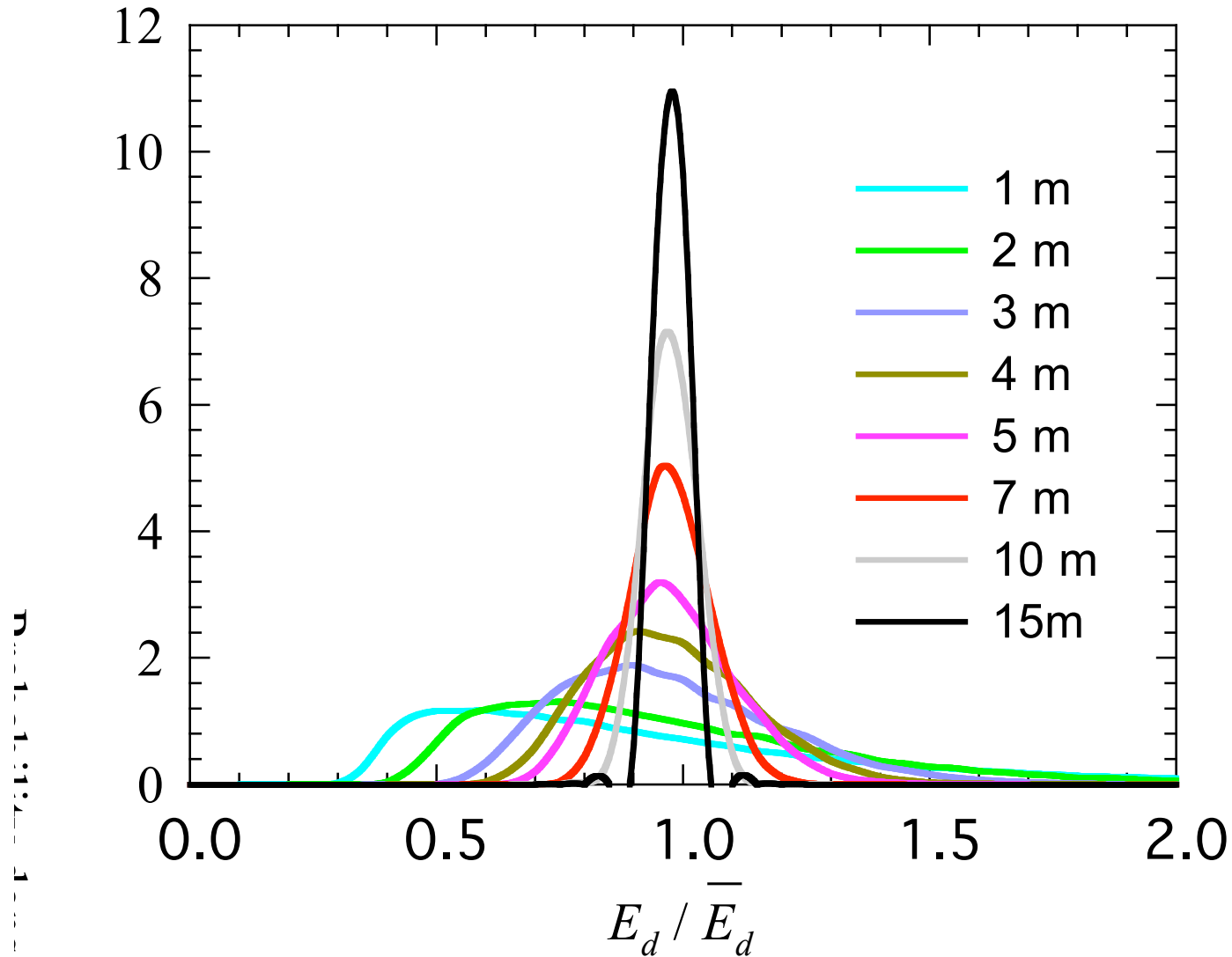
$\lambda = 532 \text{ nm}$, Clear sky, Sun $q_s = 31 - 35^\circ$, Wind = $4.1 - 6.4 \text{ m s}^{-1}$



Probability distributions of downward irradiance

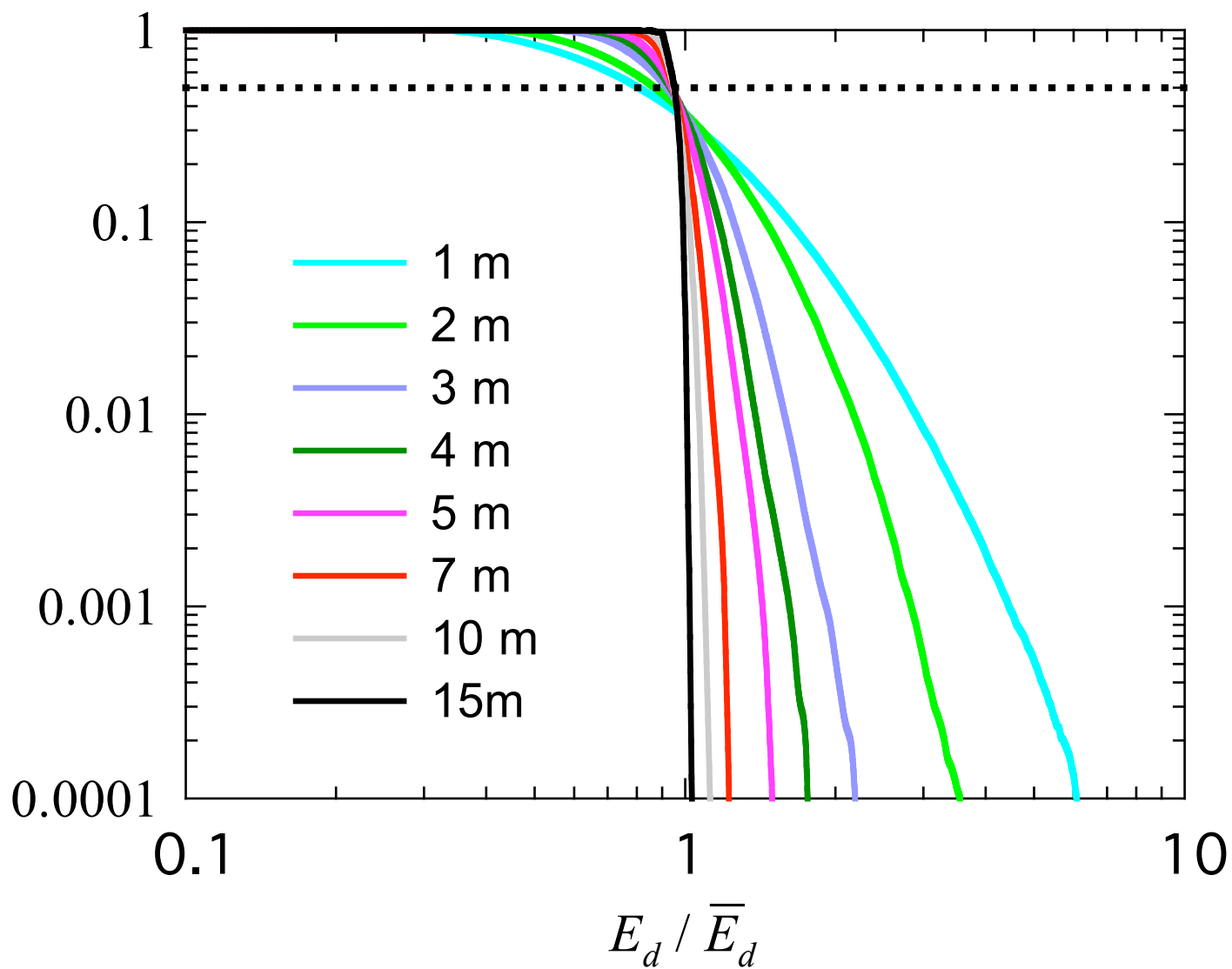
Santa Barbara Channel, 11 September 2008

$\lambda = 532 \text{ nm}$, Clear sky, Sun $q_s = 31 - 35^\circ$, Wind = $4.1 - 6.4 \text{ m s}^{-1}$



Santa Barbara Channel, 11 September 2008

$\lambda = 532 \text{ nm}$, Clear sky, Sun $q_s = 31 - 35^\circ$, Wind = $4.1 - 6.4 \text{ m s}^{-1}$



Statistical moments of downward irradiance fluctuations

Santa Barbara Channel, 11 September 2008
Clear sky, Sun $q_s = 31 - 35^\circ$, Wind = 4.1 - 6.4 m s⁻¹

